

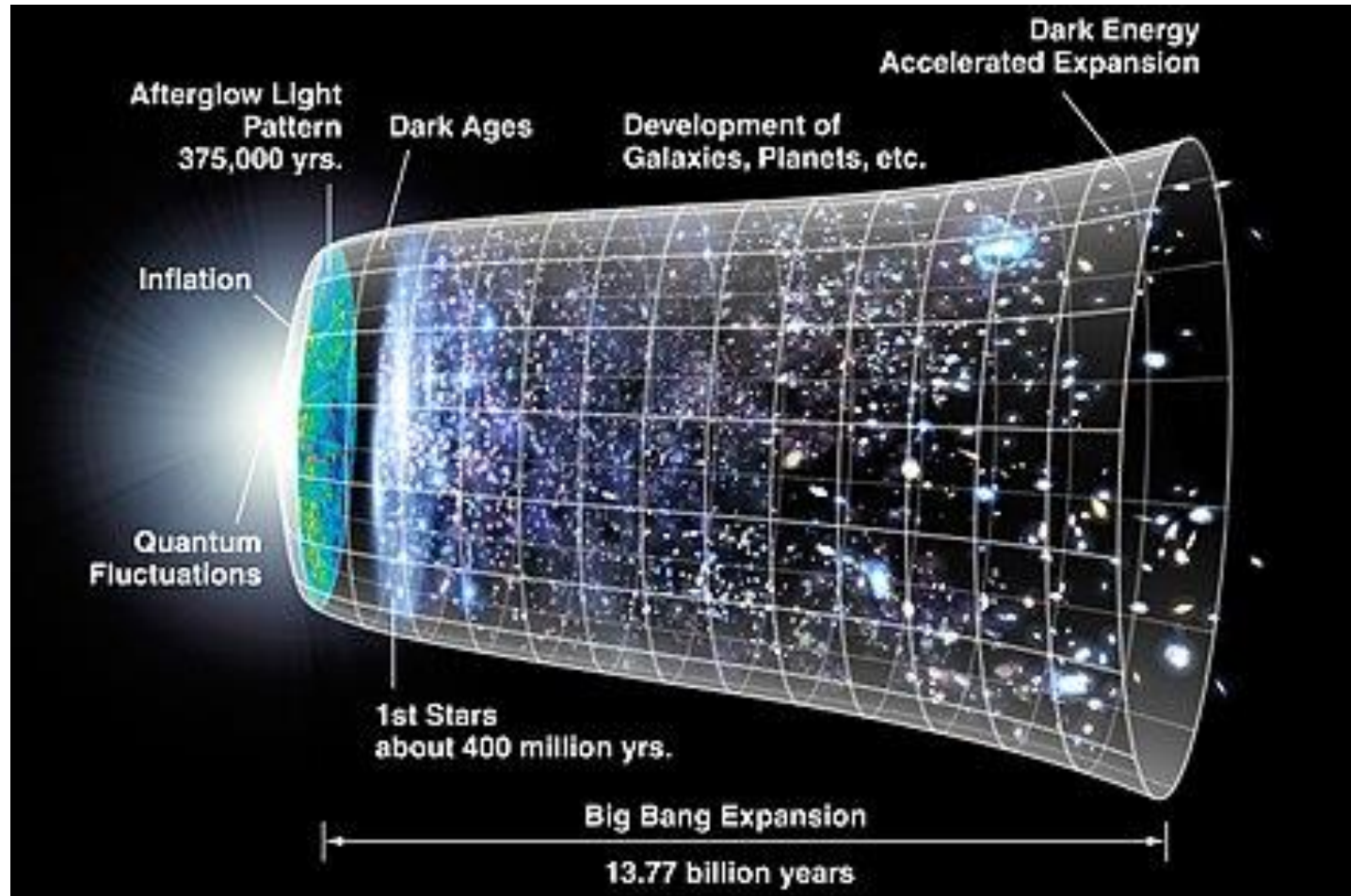
From Few-body Physics to Many-body Physics: A layman's narrative of the subject through the epochs, from Big Bang to present

M.L. Lekala
Department of Physics

Inaugural lecture, Thursday 23 August 2018 @ Thamsanqa Kambule Auditorium, Unisa Florida Campus

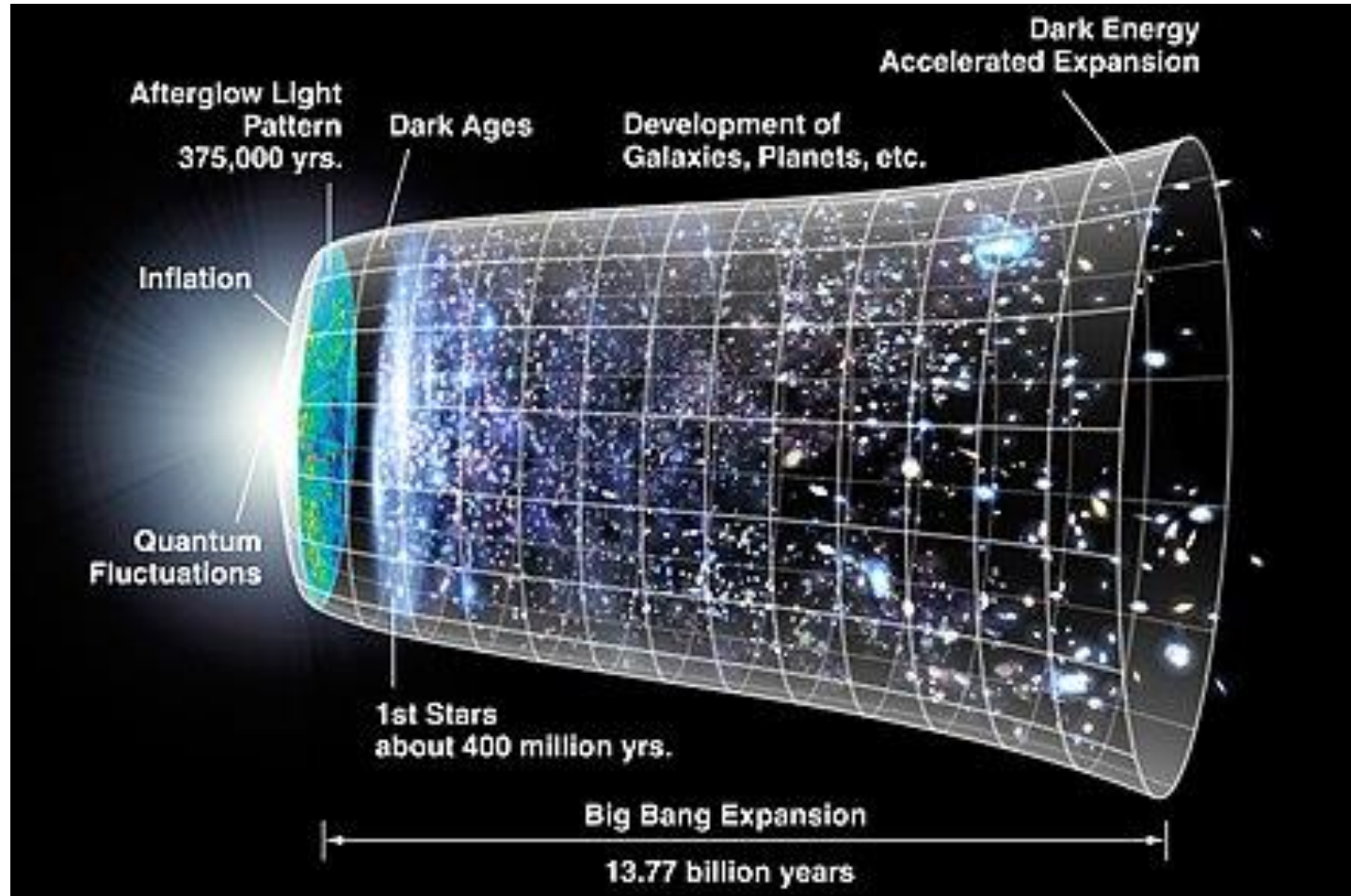


UNIVERSE



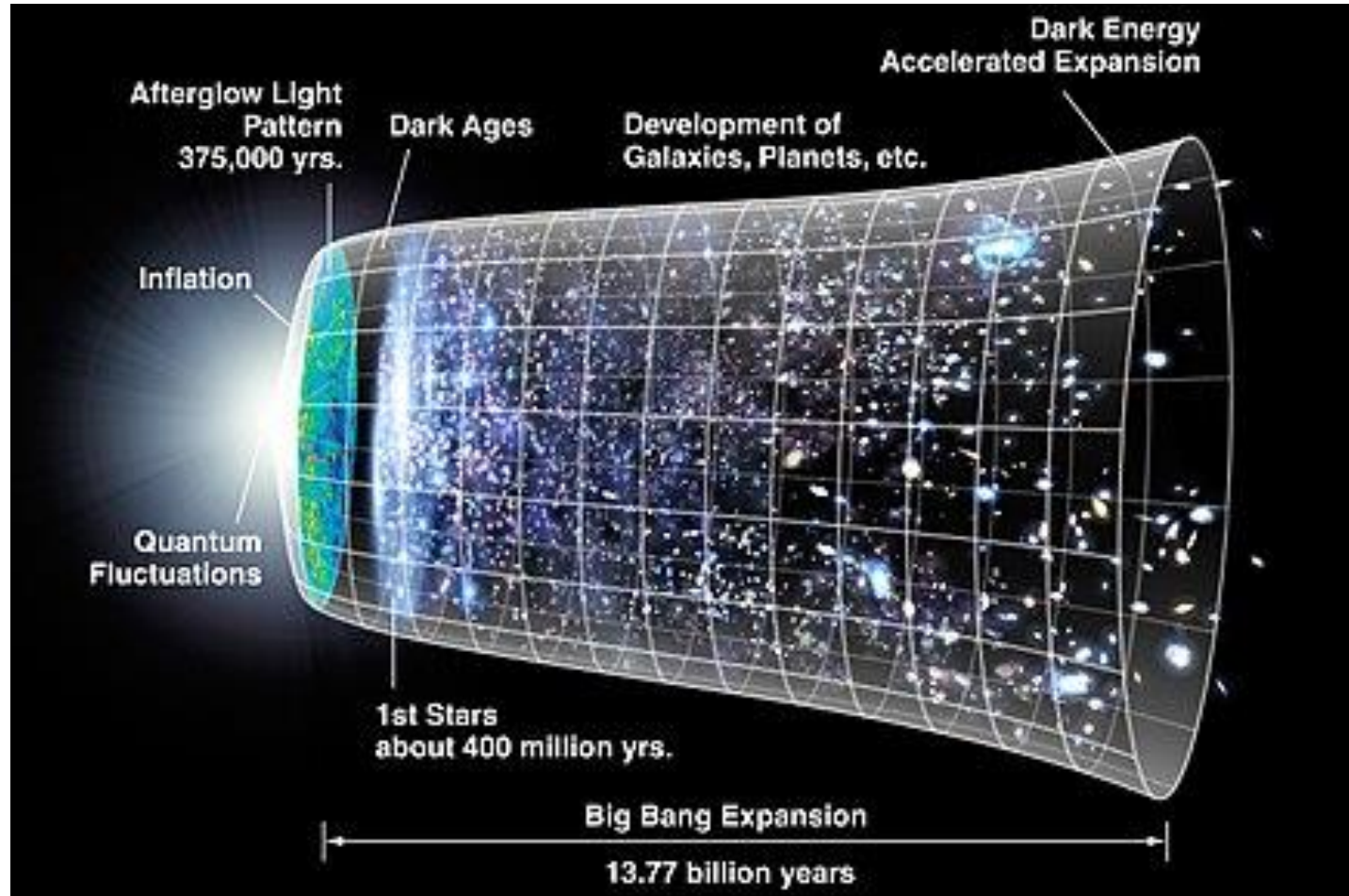
- **Very early universe**
- **Early universe**
- **Dark ages & large scale structure universe**
- **Universe as it appears now**

UNIVERSE



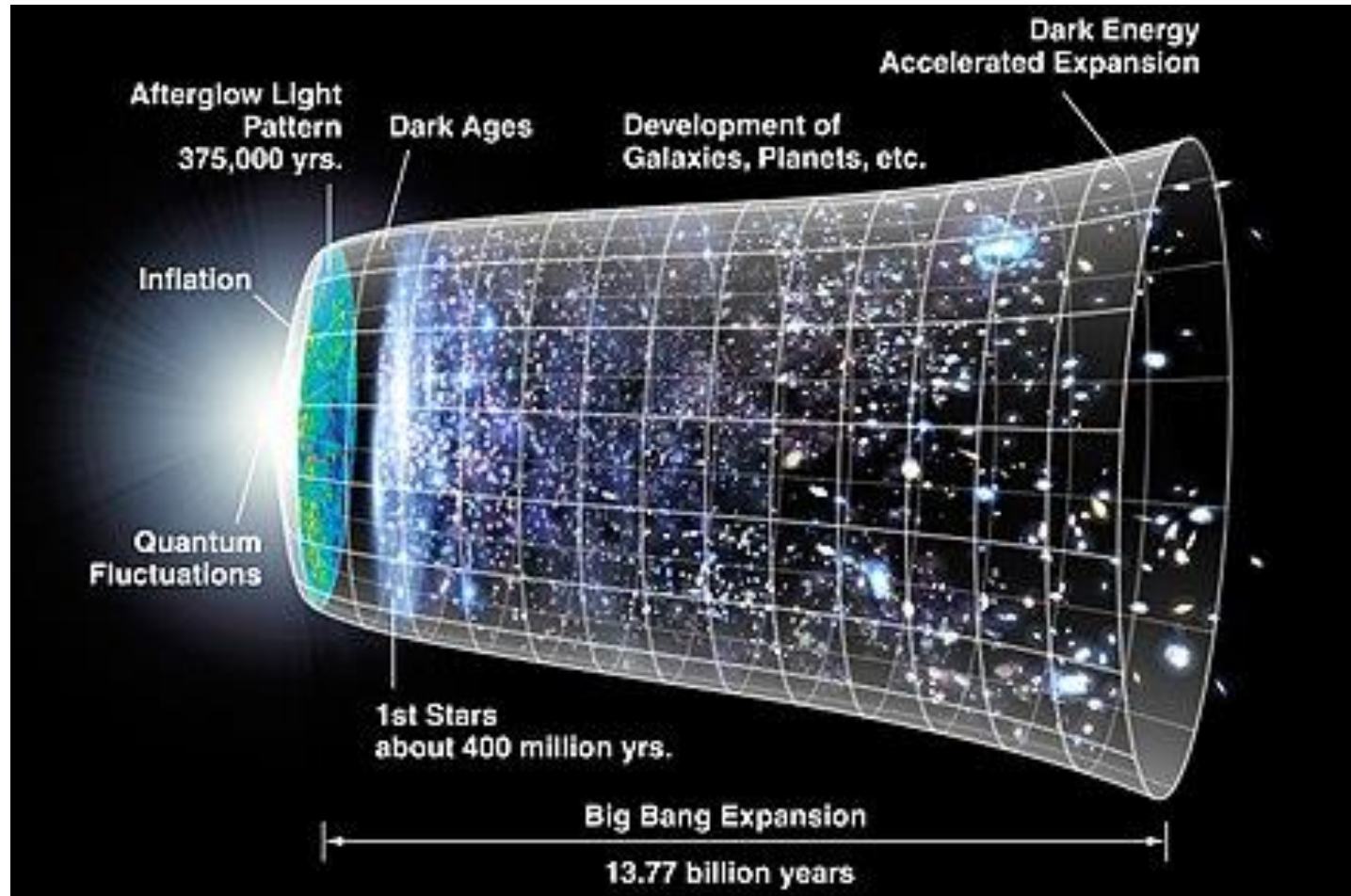
- **Very early universe**
 - ✓ **Planck epoch (10^{-12} seconds)**
 - ✓ **Grand Unification epoch**
- **Early universe**
 - ✓ **Lasted ~ 377000 years**
 - ✓ **Primordial nucleosynthesis**
- **Dark ages & large scale structure universe**
 - ✓ **377000 years → 1 billion years**
 - ✓ **Generation of galaxies**
- **Universe as it appears now**
 - ✓ **From 1 billion years → now**

UNIVERSE



- **Very early universe**
 - ✓ Planck epoch (10^{-12} seconds)
 - ✓ Grand Unification epoch
- **Early universe**
 - ✓ Lasted ~ 377000 years
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- **Dark ages & large scale structure universe**
 - ✓ 377000 years → 1 billion years
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- **Universe as it appears now**
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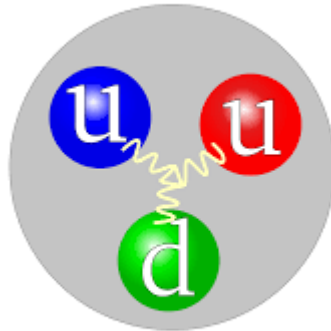
UNIVERSE



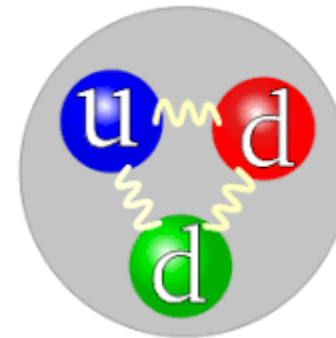
- Gravity
- Electromagnetic force
- Weak force
- Strong force

HADRONS (BARYONS + MESONS)

PROTON

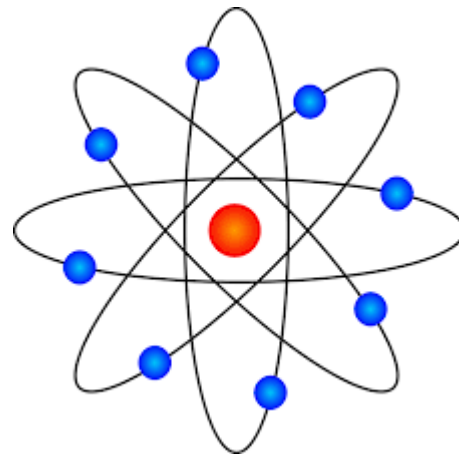


NEUTRON



NUCLEONS

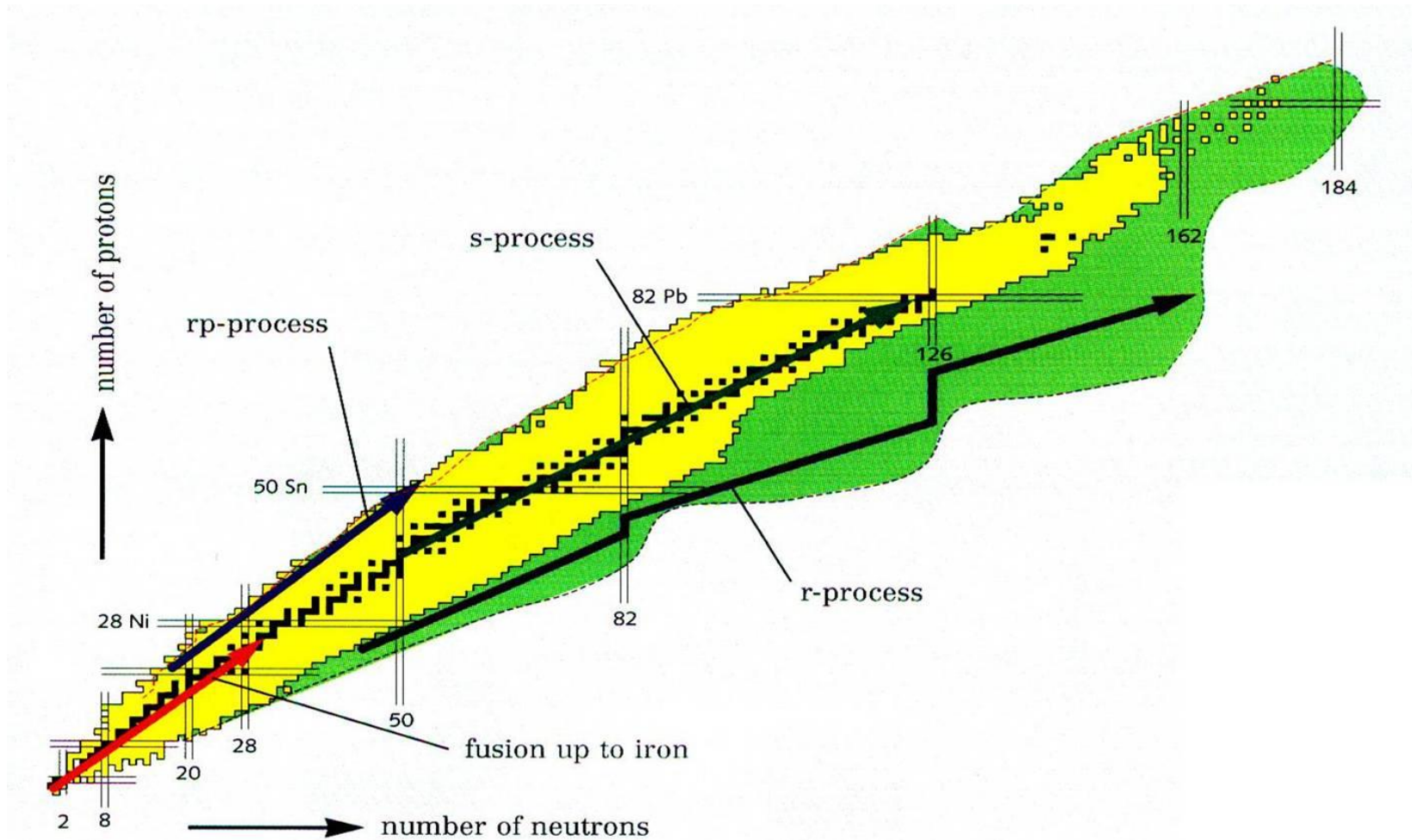
(NUCLEON-NUCLEON INTERACTION)



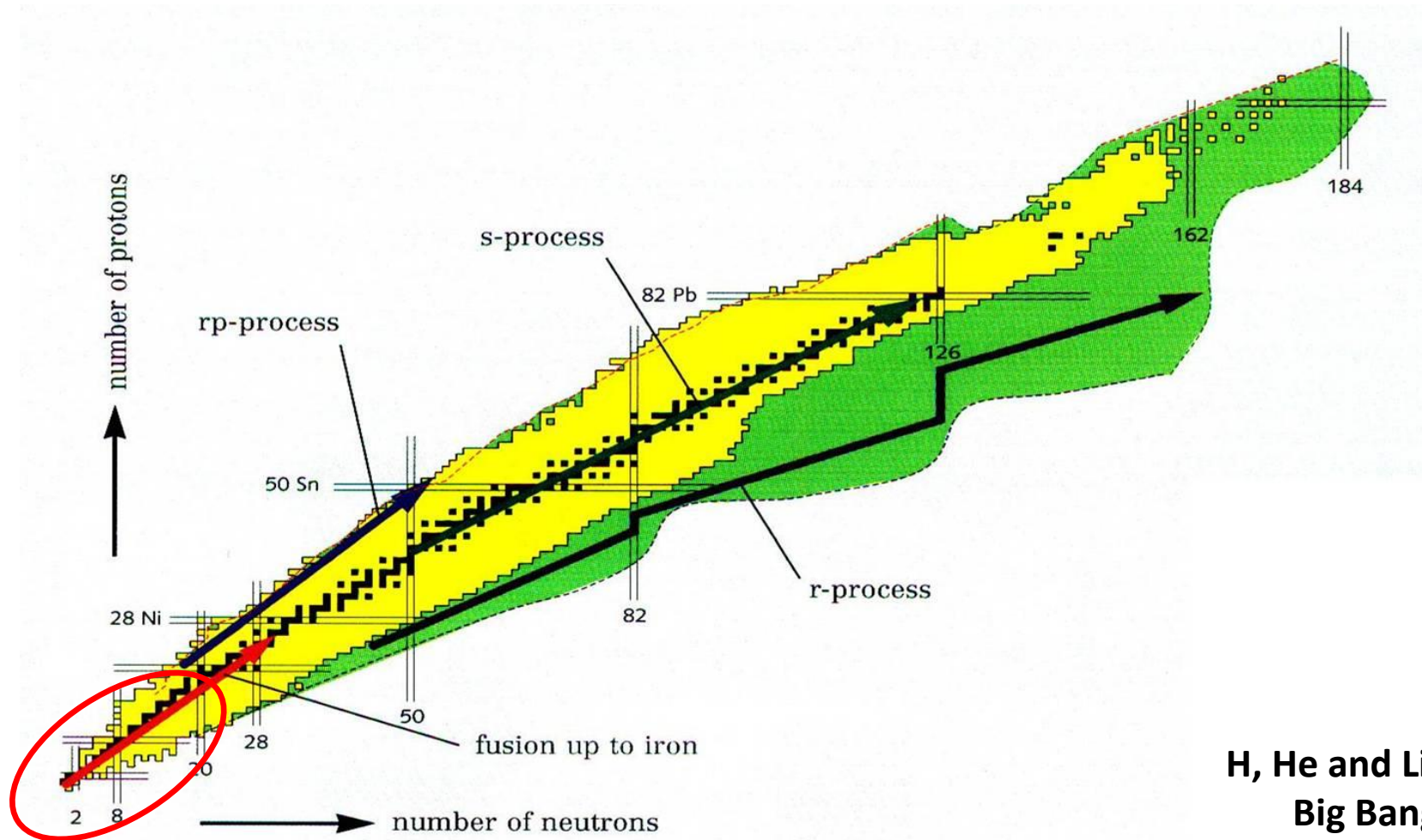
● **Nucleus = protons + neutrons**

● **Electrons**

PLAYGROUND

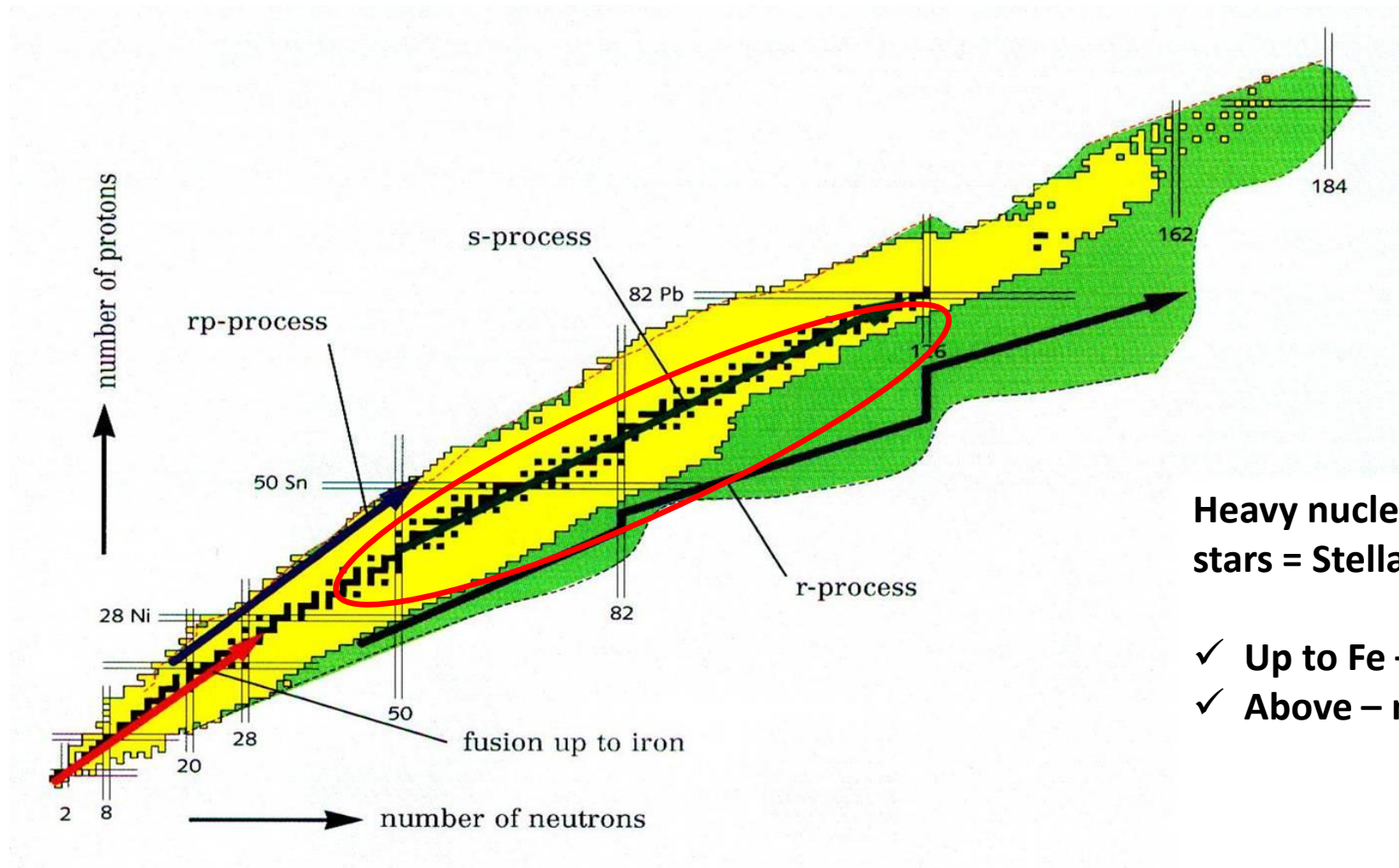


PLAYGROUND...



H, He and Li isotopes are formed during Big Bang nucleosynthesis epoch

PLAYGROUND...

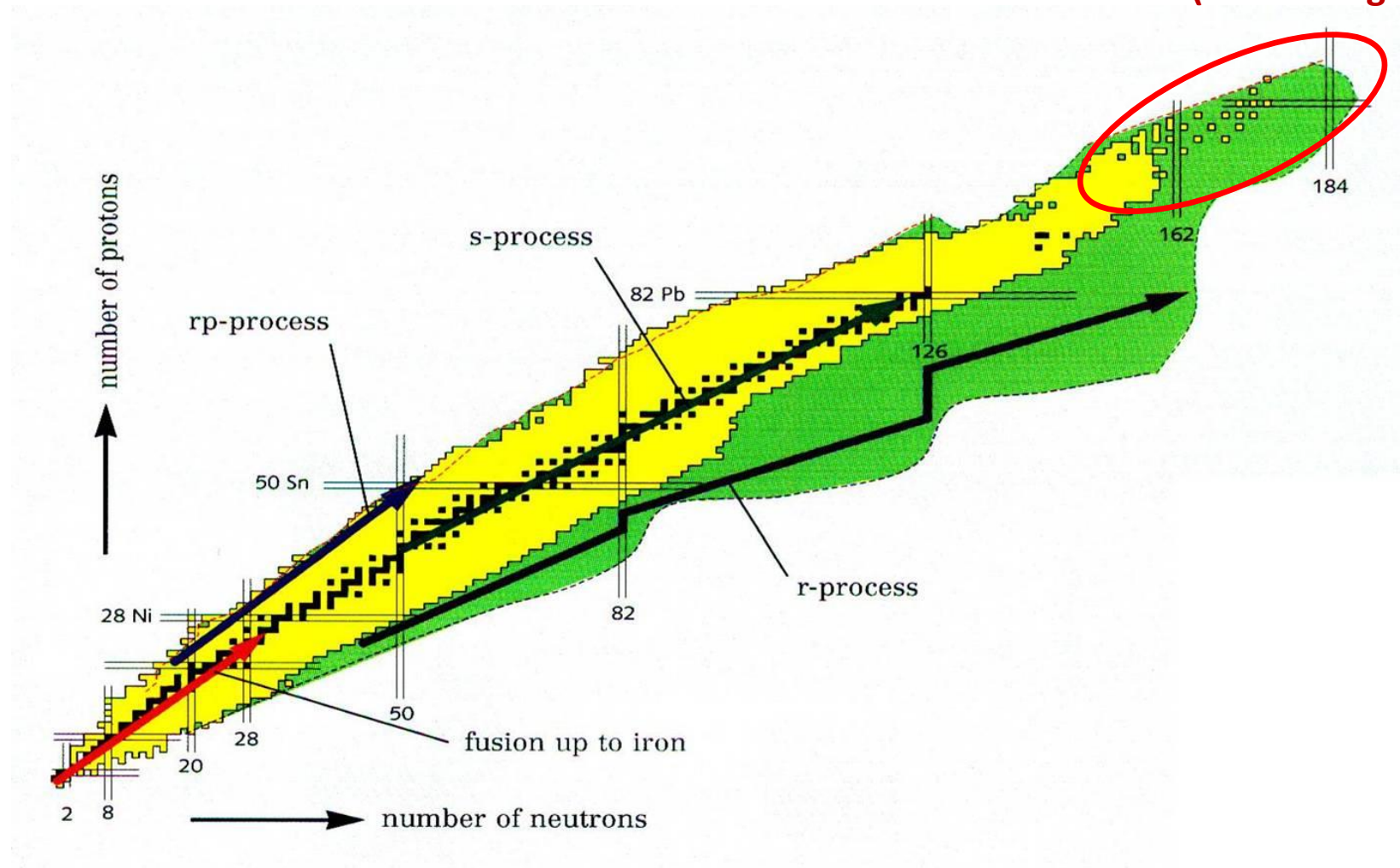


Heavy nuclei were produced in stars = Stellar nucleosynthesis epoch.

- ✓ Up to Fe – fusion,
- ✓ Above – neutron, proton capture

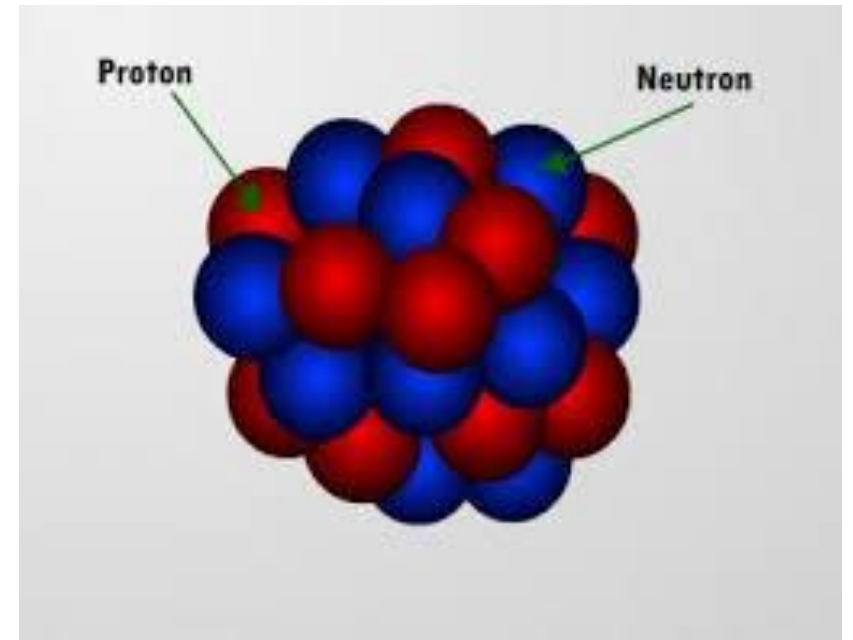
PLAYGROUND...

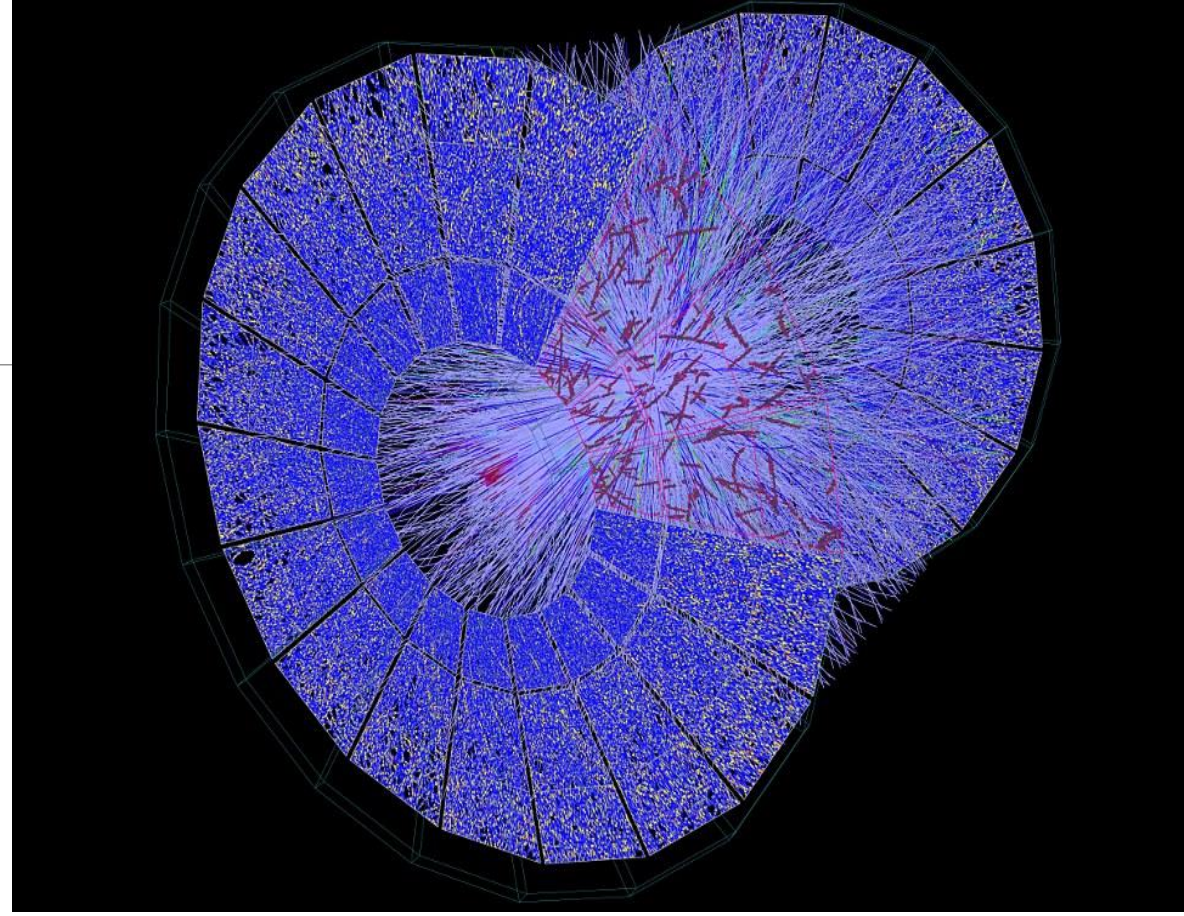
Super heavy elements, believed to be produced in neutron stars.
(terra incognita – island of stability)



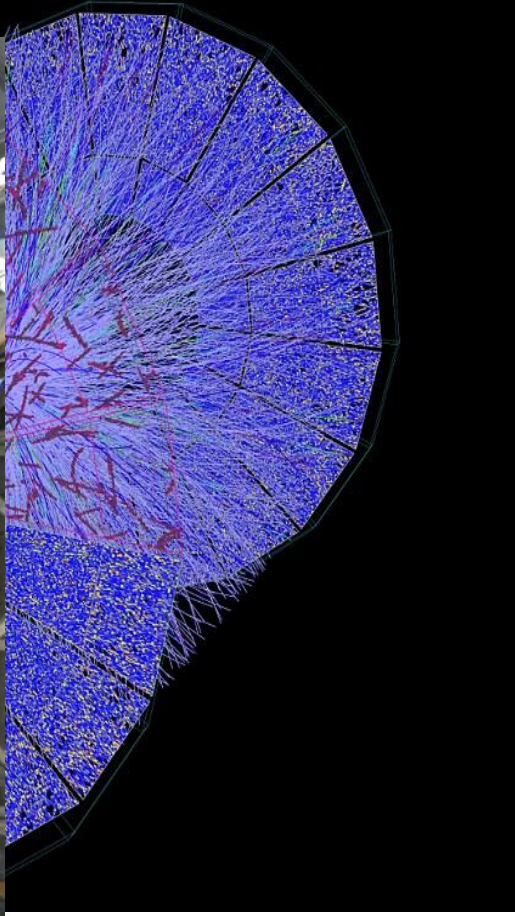
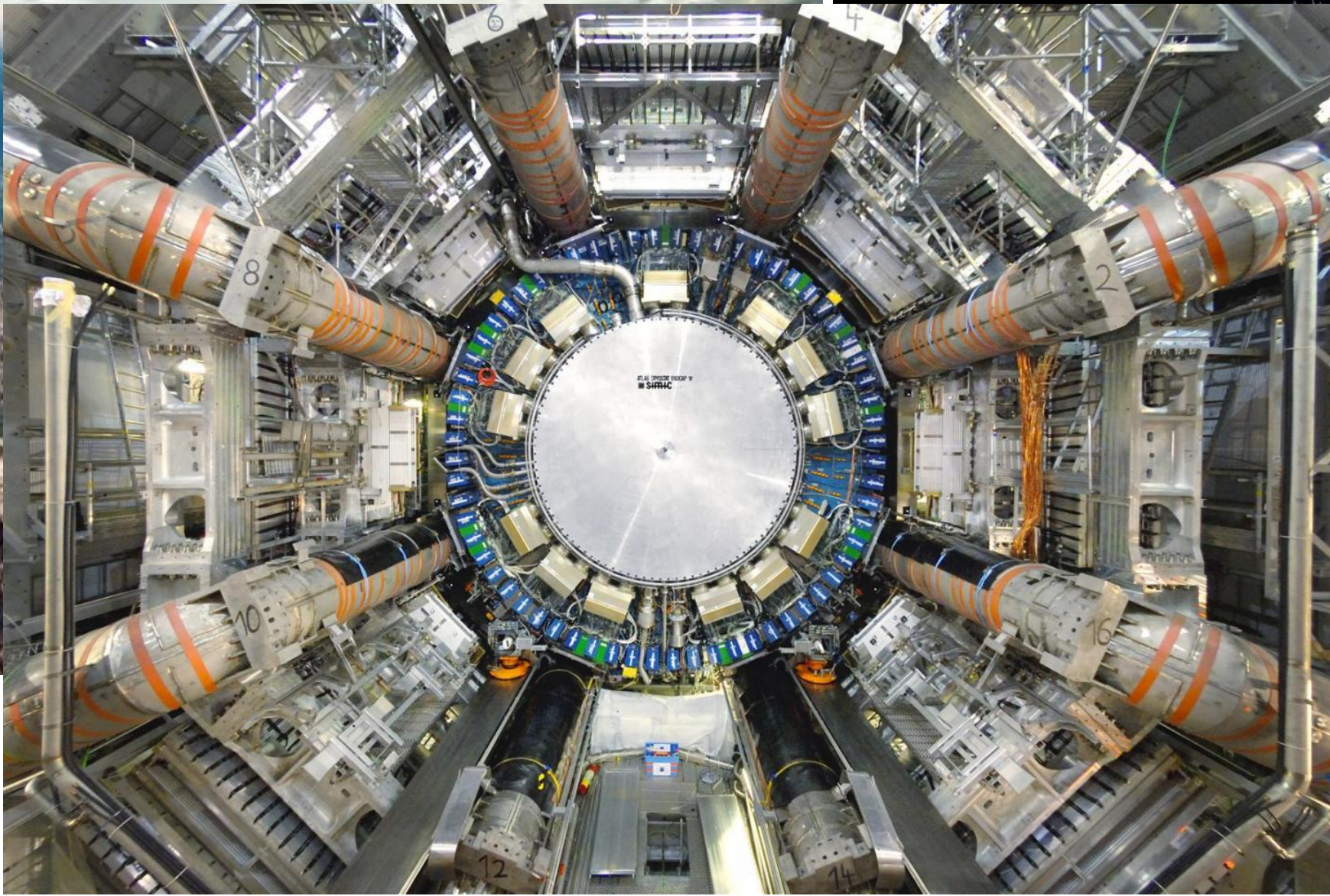
FUNDAMENTAL GOAL OF NP

- ❖ To understand the nature of the Nucleon-Nucleon interaction (strong interaction)



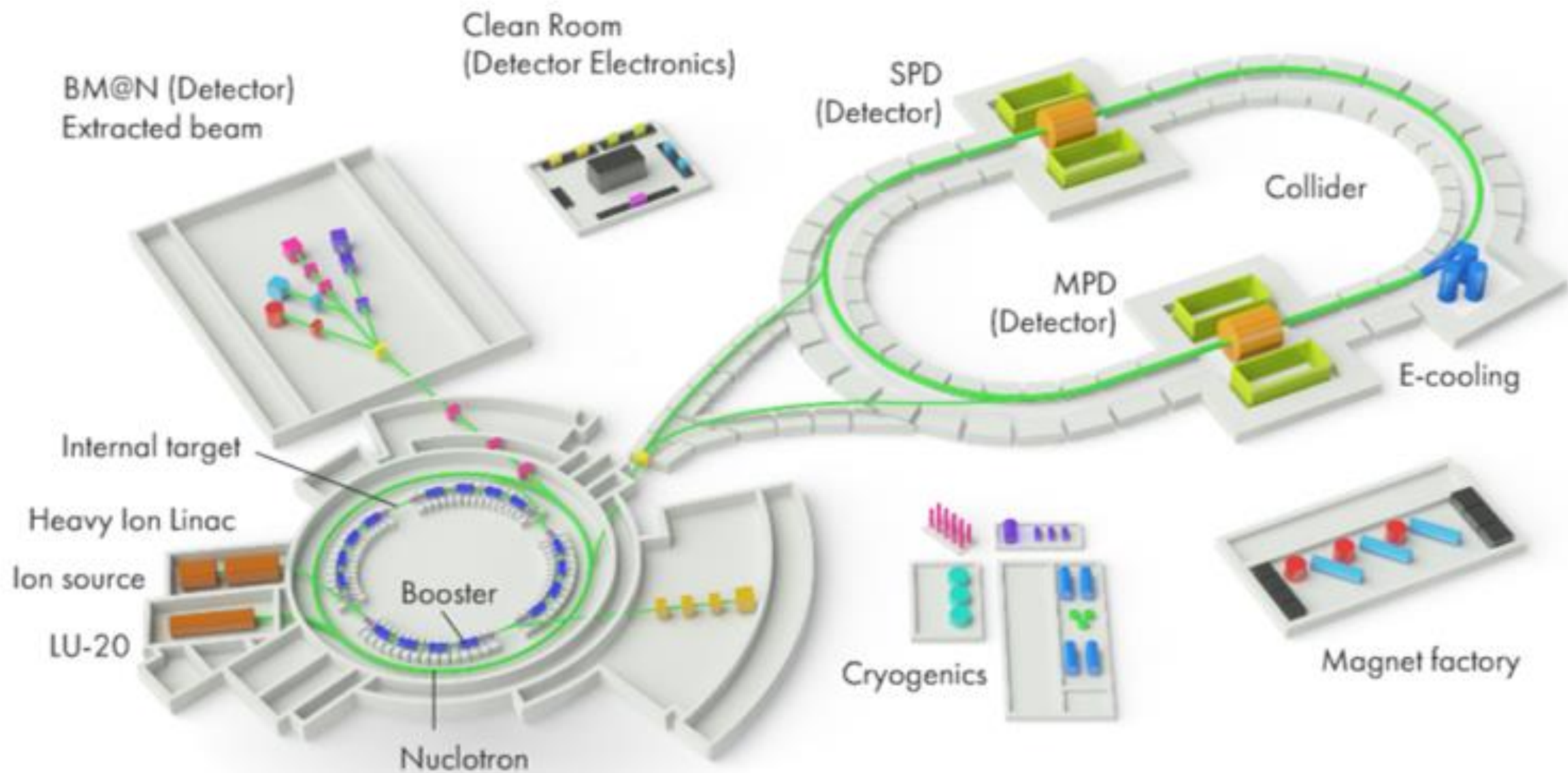


LHC (Large Hadron Collider)



Dimensions: 46 metres long, 25 metres in diameter, and weighs about 7,000 tonnes; it contains ~3000 km of cable

NICA Complex



E.g. OPEN FUNDAMENTAL QUESTIONS

- The nuclear interaction and its connection with QCD: where does the nuclear force which binds nuclei together get its main characteristics from?
- Nuclear structure and nuclear reactions: why is the nuclear landscape so rich and so complex, and where does such complexity come from?
- Hadronic physics and QCD: how can we connect the world of quarks and gluons with that of hadrons?

OPEN FUNDAMENTAL QUESTIONS...

- ❑ Relativistic heavy ion collisions: how can we understand the different nuclear phases, which lead from nuclear liquid to quark-gluon plasma?
- ❑ Nuclear astrophysics: how can we understand ultimately where the elements come from, and how does nuclear physics affect the early stage and the evolution of our Universe?

OVERVIEW

- What is Few-body Physics?
- Fundamental equations of few-body physics
- Applications of Few-body Physics (snapshot of our work)
- Summary / Conclusions



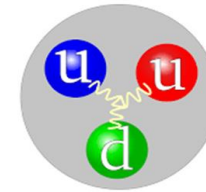
WHAT IS FEW-BODY PHYSICS

Definition

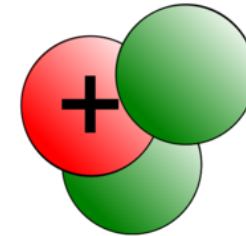
□ A sufficiently simple & isolated system that permits:

- ✓ A complete experiment on the system.
- ✓ This definition depends on the current state of the art in experimental physics, theoretical physics, and computational physics.

E.g.



Proton



Triton



${}^3\text{He}$



${}^4\text{He}$



${}^6\text{He}$

WHAT IS FEW-BODY PHYSICS

Definition:

A sufficiently simple & isolated system that permits:

- ✓ A complete experiment on the system.
- ✓ This definition depends on the current state of the art in experimental physics, theoretical physics, and computational physics.

Theories should satisfy:

Quantum theory

Relativistically invariant

Should be possible to do ab-initio computations

Should satisfy cluster properties
(link between few-body and many-body problems)

FUNDAMENTAL EQUATIONS



(1887 – 1961)

$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 + V \right] \Psi = -i\hbar \frac{\partial \Psi}{\partial t}$$



$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 + V \right] \Psi = E\Psi$$

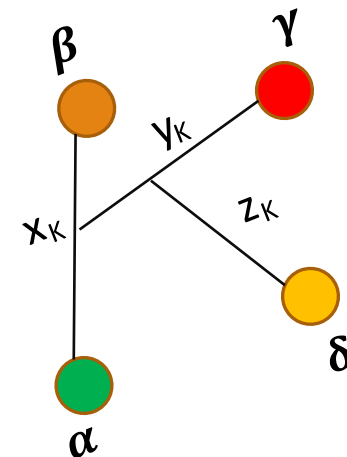
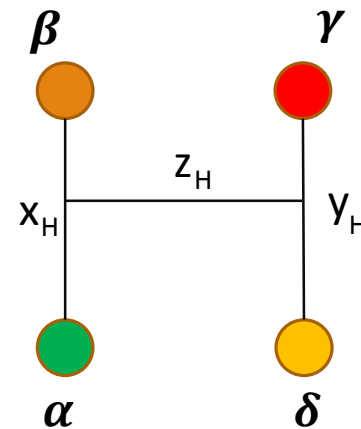
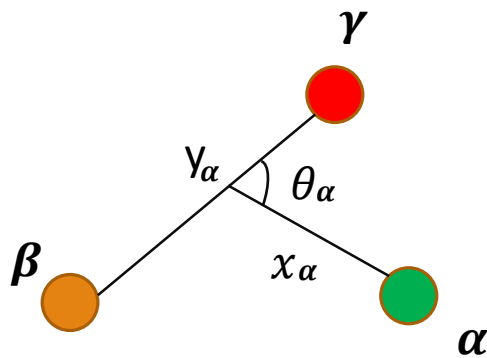
2B (tractable). 3B (intractable)

FUNDAMENTAL EQUATIONS

3 BODY SYSTEM

4 BODY SYSTEM

(JACOBI COORDINATES)



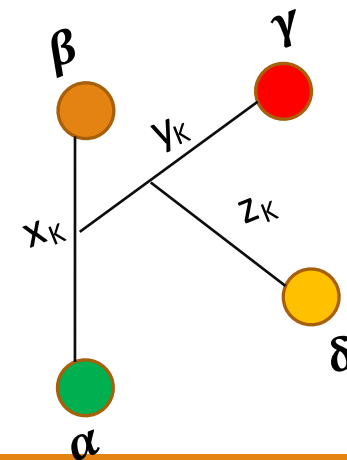
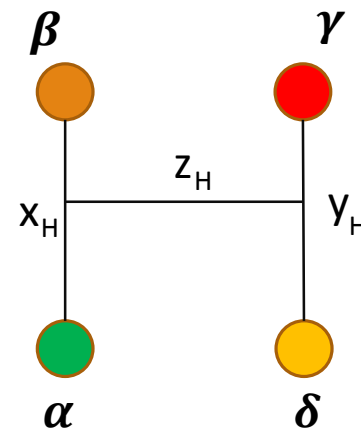
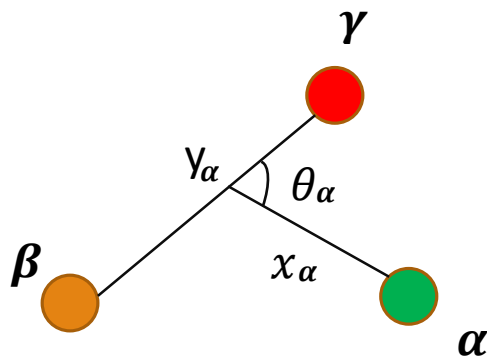
FUNDAMENTAL EQUATIONS

$$\Psi_{3B}(x, y) = \sum_{i=\alpha, \beta, \gamma} \psi_i(x_i, y_i)$$

$$(H_0 + V_\alpha(r_\alpha) - E_{3B})\psi_\alpha = -V_\alpha(\psi_\beta + \psi_\gamma)$$



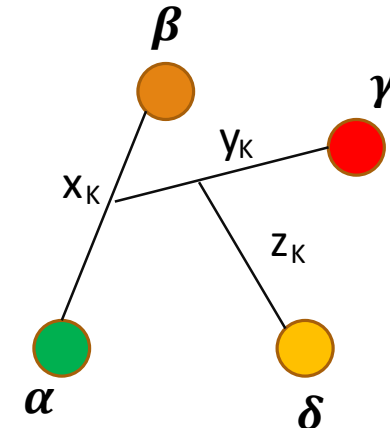
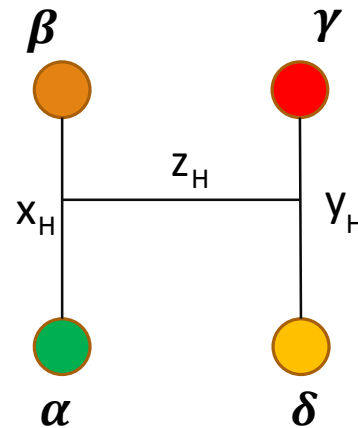
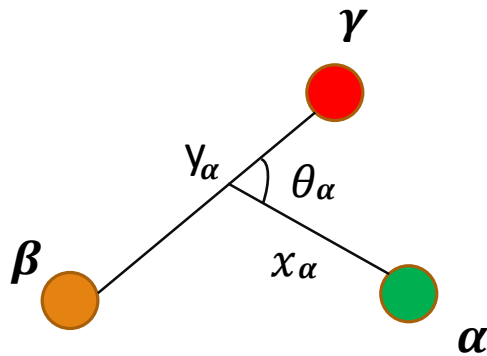
(1934 – 2017)



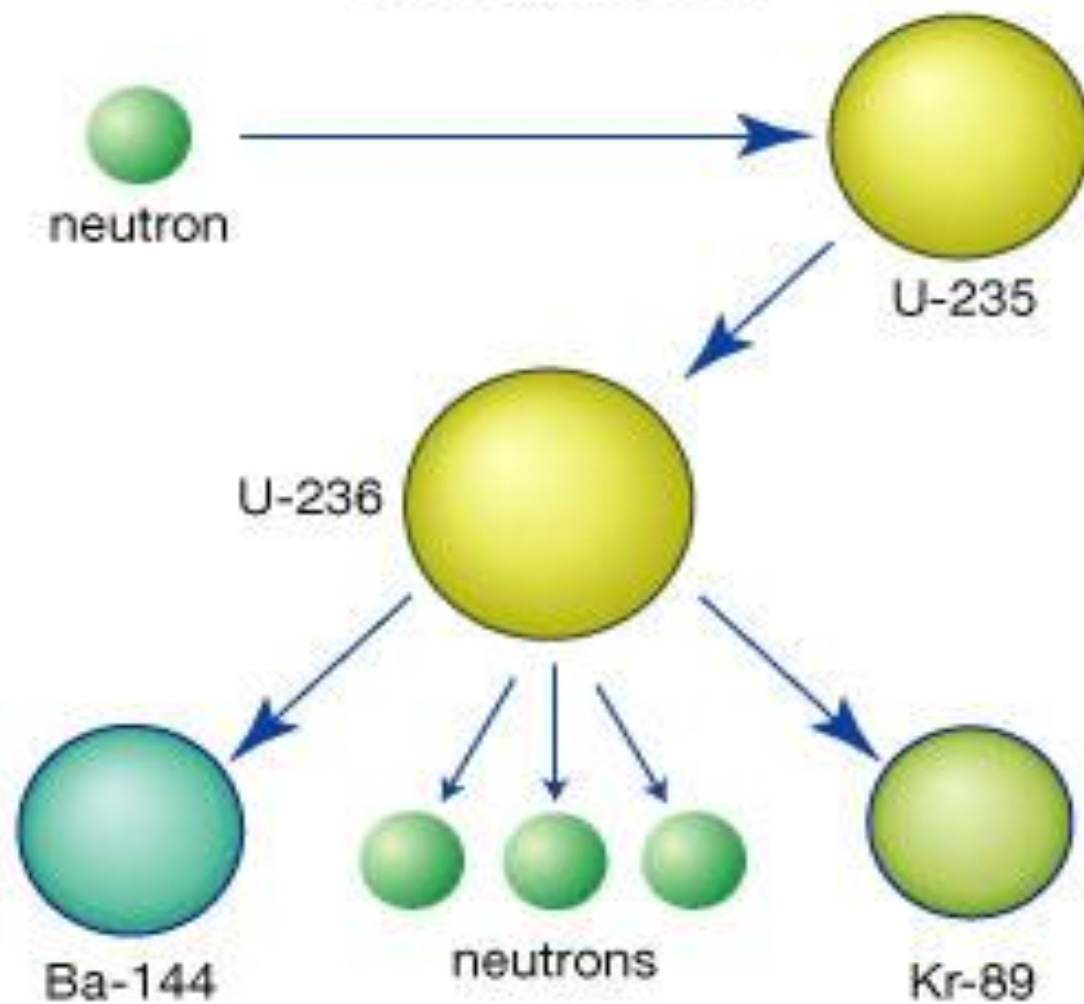
FUNDAMENTAL EQUATIONS

$$(H_0 + V_{\alpha\beta} - E_{4B})\psi_{(\alpha\beta)(\gamma\delta)} = -V_{\alpha\beta} (\psi_{\gamma\delta,\alpha}^{\beta} + \psi_{\gamma\delta,\beta}^{\alpha} + \psi_{(\gamma\alpha)(\alpha\beta)})$$

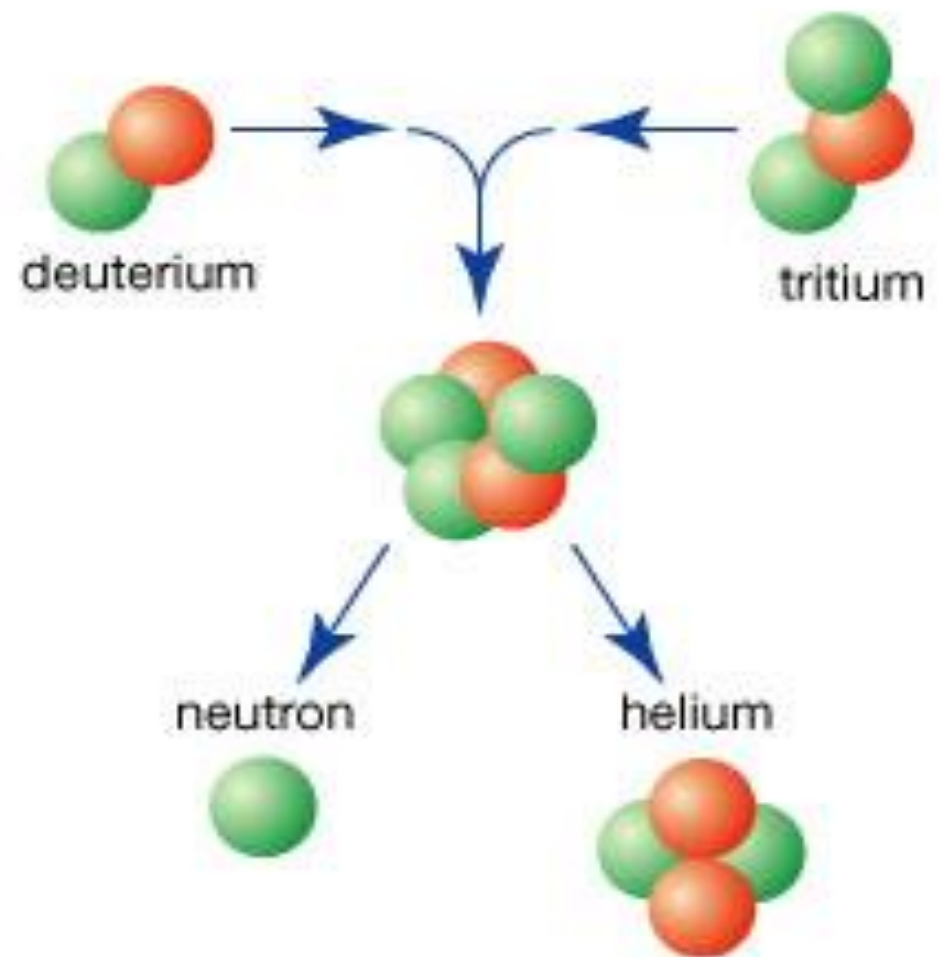
$$(H_0 + V_{\alpha\beta} - E_{4B})\psi_{\alpha,\beta,\gamma}^{\delta} = -V_{\alpha\beta} (\psi_{\alpha\gamma,\beta}^{\delta} + \psi_{\alpha\gamma,\delta}^{\beta} + \psi_{(\alpha\gamma)(\beta\delta)} + \psi_{\beta\gamma,\alpha}^{\delta} + \psi_{\beta\gamma,\delta}^{\alpha} + \psi_{(\beta\gamma)(\alpha\delta)})$$



nuclear fission



nuclear fusion



NUCLEAR REACTIONS



Available online at www.sciencedirect.com

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Nuclear Physics A 969 (2018) 60–67

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NUCLEAR
PHYSICS A

For energies above the Coulomb barrier

- breakup cross sections are increased by inclusion of resonances ($\sim 10\%$),
- fusion cross sections are reduced ($\sim 4\%$).

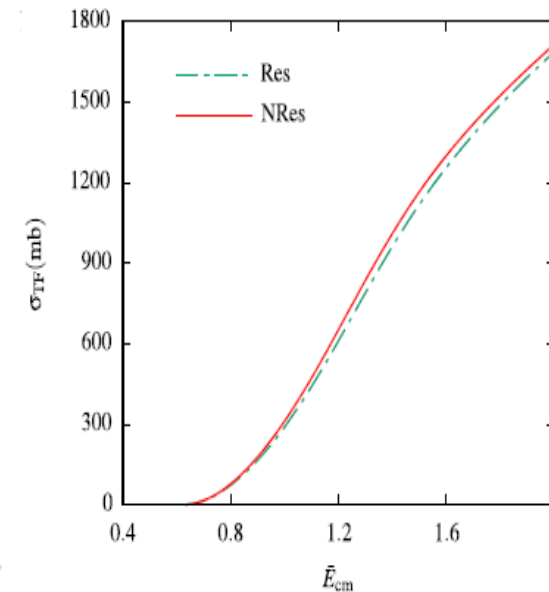
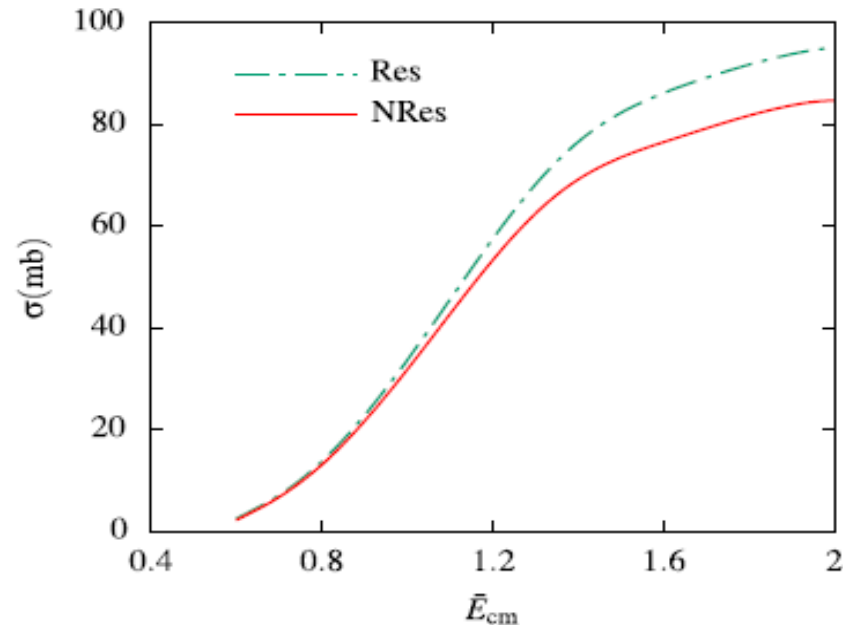
Role of projectile resonances on breakup and fusion cross sections in the ${}^6\text{Li} + {}^{144}\text{Sm}$ reaction

B. Mukeru *, G.J. Rampho, M.L. Lekala

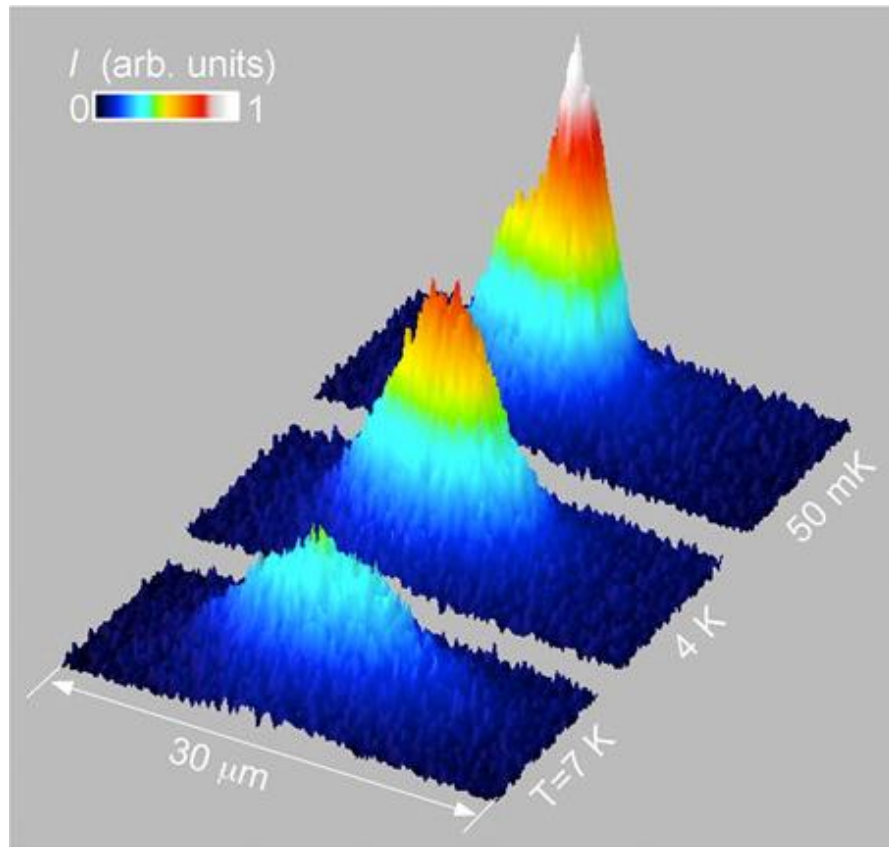
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BOSE-EINSTEIN CONDENSATION



JILA Experiment (1995)

Path to Bose-Einstein Condensation

- A dilute gas of alkali atoms
- Extremely good vacuum & evaporative cooling
- Tune the interaction strength between particles by using Feshbach resonances
- Shape the trap

Theories for Bose-Einstein Condensates

- The many-body Schrodinger equation [MBSE] (very difficult to solve)
- Most approximate method to solve MBSE: Gross-Pitaevskii theory

MOLECULAR SYSTEMS

Few-Body Syst
DOI 10.1007/s00601-013-0705-2

S. A. Sofianos · T. K. Das · B. Chakrabarti · M. L. Lekala ·
R. M. Adam · G. J. Rampho

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Spectral analysis of molecular resonances in erbium isotopes: Are they close to semi-Poisson?

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and G. J. RAMPHO⁴

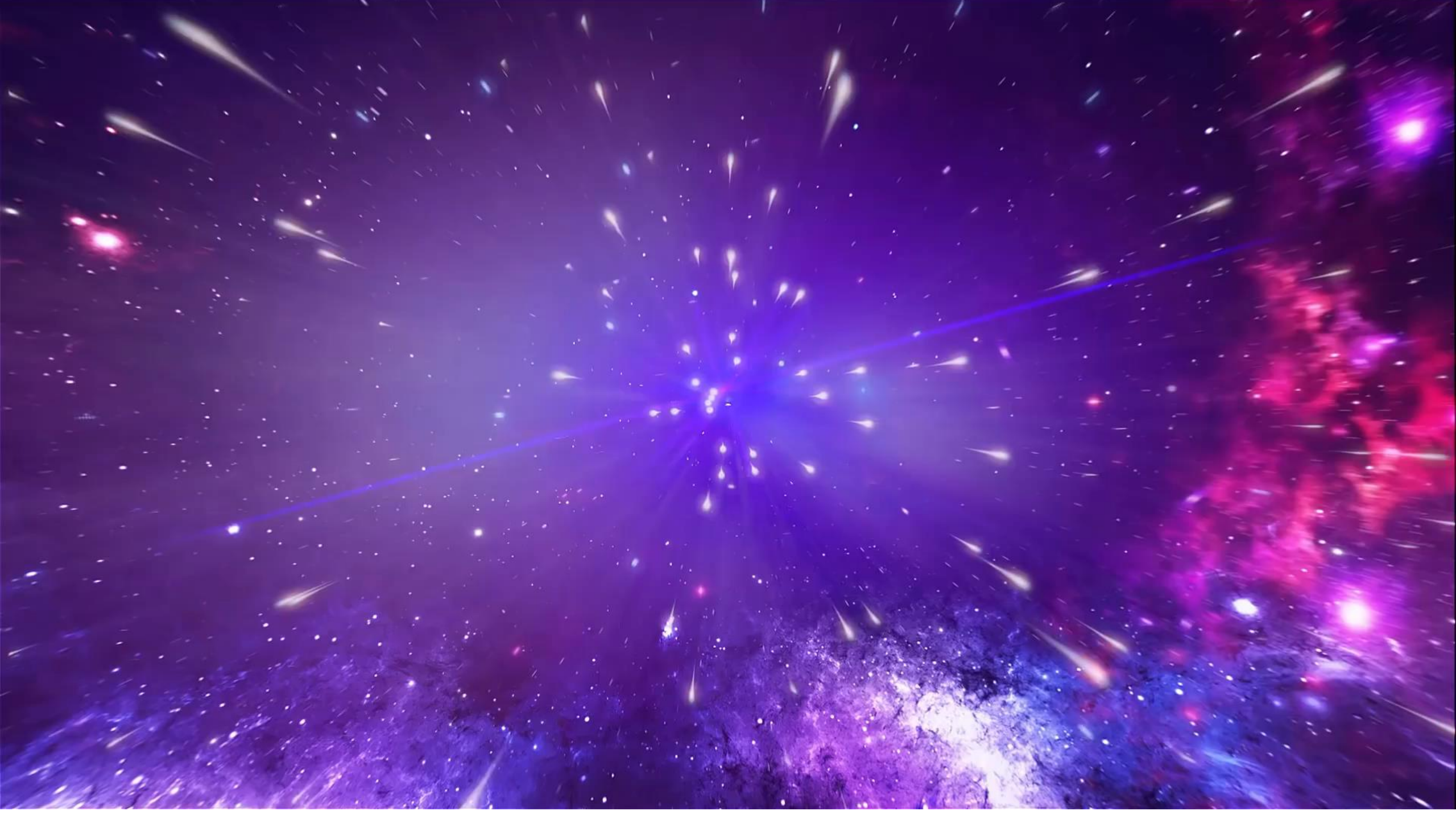
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SUMMARY

- ❑ The field of Few-Body Physics is a very active and exciting field.
- ❑ Few-body physics techniques are applicable to studying systems and phenomenon in a wide range of fields, from few-body systems to many-body systems such as e.g. BEC, etc.
- ❑ The FB techniques can be applied to reveal the rich structure of the fundamental components of matter. Thus the understanding of the Universe.



Veni, vidi, vici.
(I came, I saw, I conquered.)

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EXTRAS